Extra Dimensions

For explanation of terms used and discussion of significant model dependence of following limits, see the "Extra Dimensions" review. Footnotes describe originally quoted limit. δ indicates the number of extra dimensions.

Limits not encoded here are summarized in the "Extra Dimensions" review, where the latest unpublished results are also described.

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CONTENTS:

Limits on R from Deviations in Gravitational Force Law Limits on R from On-Shell Production of Gravitons: $\delta=2$ Mass Limits on M_{TT} Limits on $1/R=M_C$

Limits on Kaluza-Klein Gravitons in Warped Extra Dimensions

Limits on R from Deviations in Gravitational Force Law

This section includes limits on the size of extra dimensions from deviations in the Newtonian $(1/r^2)$ gravitational force law at short distances. Deviations are parametrized by a gravitational potential of the form $V=-(G\ m\ m'/r)\ [1+\alpha\ \exp(-r/R)]$. For δ toroidal extra dimensions of equal size, $\alpha=8\delta/3$. Quoted bounds are for $\delta=2$ unless otherwise noted.

VALUE (μm)	CL%	DOCUMENT ID		COMMENT
< 30	95	¹ KAPNER	07	Torsion pendulum
• • • We do not use the	following	g data for averages	s, fits,	limits, etc. • • •
		² BEZERRA	11	Torsion oscillator
		³ SUSHKOV	11	Torsion pendulum
		⁴ BEZERRA	10	Microcantilever
		⁵ MASUDA	09	Torsion pendulum
		⁶ GERACI	80	Microcantilever
		⁷ TRENKEL	80	Newton's constant
		⁸ DECCA	07A	Torsion oscillator
< 47	95	⁹ TU	07	Torsion pendulum
		¹⁰ SMULLIN	05	Microcantilever
<130	95	¹¹ HOYLE	04	Torsion pendulum
		12 CHIAVERINI	03	Microcantilever
$\lesssim 200$	95	¹³ LONG	03	Microcantilever
<190	95	14 HOYLE	01	Torsion pendulum
		¹⁵ HOSKINS	85	Torsion pendulum

 1 KAPNER 07 search for new forces, probing a range of $\alpha\simeq 10^{-3}$ –10 5 and length scales $R\simeq 10$ –1000 $\mu\mathrm{m}$. For $\delta=1$ the bound on R is 44 $\mu\mathrm{m}$. For $\delta=2$, the bound is expressed in terms of M_{*} , here translated to a bound on the radius. See their Fig. 6 for details on the bound.

 2 BEZERRA 1 obtain constraints on non-Newtonian forces with strengths $^{101} \lesssim |\alpha| \lesssim 10^{18}$ and length scales R=30– 1 260 nm. See their Fig. 2 for more details. These constraints do not place limits on the size of extra flat dimensions.

 3 SUSHKOV 11 obtain improved limits on non-Newtonian forces with strengths $10^7 \lesssim |\alpha| \lesssim 10^{11}$ and length scales 0.4 $\mu \rm m < R < 4~\mu m$ (95% CL). See their Fig. 2. These bounds do not place limits on the size of extra flat dimensions. However, a model dependent bound of $M_* > 70$ TeV is obtained assuming gauge bosons that couple to baryon number also propagate in (4 + δ) dimensions.

 4 BEZERRA 10 obtain improved constraints on non-Newtonian forces with strengths $10^{19}\lesssim |\alpha|\lesssim 10^{29}$ and length scales R=1.6–14 nm (95% CL). See their Fig. 1. This bound does not place limits on the size of extra flat dimensions.

 5 MASUDA 09 obtain improved constraints on non-Newtonian forces with strengths $10^9 \lesssim |\alpha| \lesssim 10^{11}$ and length scales R=1.0–2.9 μm (95% CL). See their Fig. 3. This bound does not place limits on the size of extra flat dimensions.

 6 GERACI 08 obtain improved constraints on non-Newtonian forces with strengths $|\alpha|>14,000$ and length scales R=5–15 $\mu{\rm m}.$ See their Fig. 9. This bound does not place limits on the size of extra flat dimensions.

 7 TRENKEL 08 uses two independent measurements of Newton's constant $\it G$ to constrain new forces with strength $|\alpha| \simeq 10^{-4}$ and length scales $\it R = 0.02$ –1 m. See their Fig. 1. This bound does not place limits on the size of extra flat dimensions.

NODE=S071

NODE=S071

NODE=S071CNT NODE=S071CNT

NODE=S071CNT

NODE=S071DGF NODE=S071DGF

NODE=S071DGF

NODE=S071DGF;LINKAGE=KA

NODE=S071DGF;LINKAGE=BZ

NODE=S071DGF;LINKAGE=SU

NODE=S071DGF;LINKAGE=BE

NODE=S071DGF;LINKAGE=MA

NODE=S071DGF;LINKAGE=GE

NODE=S071DGF;LINKAGE=TR

 8 DECCA 07A search for new forces and obtain bounds in the region with strengths $|lpha|~\simeq$ 10^{13} – 10^{18} and length scales R=20–86 nm. See their Fig. 6. This bound does not place limits on the size of extra flat dimensions.

⁹ TU 07 search for new forces probing a range of $|\alpha| \simeq 10^{-1}$ – 10^5 and length scales $R \simeq 20$ – $1000~\mu m$. For $\delta = 1$ the bound on R is 53 μm . See their Fig. 3 for details on the

 $10\,\mathrm{SMULLIN}$ 05 search for new forces, and obtain bounds in the region with strengths $\alpha \simeq 10^3 - 10^8$ and length scales $R = 6 - 20~\mu m$. See their Figs. 1 and 16 for details on the bound. This work does not place limits on the size of extra flat dimensions.

 $^{11}\, \rm HOYLE~04$ search for new forces, probing α down to 10^{-2} and distances down to $10\mu m$. Quoted bound on R is for $\delta=2$. For $\delta=1$, bound goes to 160 μm . See their Fig. 34 for details on the bound.

12 CHIAVERINI 03 search for new forces, probing α above 10⁴ and λ down to 3 μ m, finding no signal. See their Fig. 4 for details on the bound. This bound does not place limits on the size of extra flat dimensions.

 13 LONG 03 search for new forces, probing lpha down to 3, and distances down to about $10\mu m$. See their Fig. 4 for details on the bound.

 14 HOYLE 01 search for new forces, probing α down to 10^{-2} and distances down to $20\mu m$. See their Fig. 4 for details on the bound. The quoted bound is for $\alpha \geq 3$.

 15 HOSKINS 85 search for new forces, probing distances down to 4 mm. See their Fig. 13 for details on the bound. This bound does not place limits on the size of extra flat dimensions

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NODE=S071DGF;LINKAGE=TU

NODE=S071DGF;LINKAGE=SM

NODE=S071DGF;LINKAGE=HO

NODE=S071DGF;LINKAGE=C

NODE=S071DGF:LINKAGE=L

NODE=S071DGF;LINKAGE=HL

NODE=S071DGF;LINKAGE=HK

NODE=S071OS6 NODE=S071OS6

Limits on R from On-Shell Production of Gravitons: $\delta = 2$

This section includes limits on on-shell production of gravitons in collider and astrophysical processes. Bounds quoted are on R, the assumed common radius of the flat extra dimensions, for $\delta=2$ extra dimensions. Studies often quote bounds in terms of derived parameter; experiments are actually sensitive to the masses of the KK gravitons: $m_{\vec{n}} = |\vec{n}|/R$. See the Review on "Extra Dimensions" for details. Bounds are given in μ m for $\delta = 2$.

$VALUE~(\mu m)$	CL%	DOCUMENT ID		TECN	COMMENT
< 23	95	¹ CHATRCHYAN	12 AP	CMS	$pp \rightarrow jG$
\bullet \bullet We do not use the	following	g data for averages	, fits, l	limits, e	tc. • • •
< 127	95	² AAD	13 C	ATLS	$pp \rightarrow \gamma G$
< 34.4	95	³ AAD	13 D	ATLS	$pp \rightarrow jj$
< 0.0087	95	⁴ AJELLO	12	FRMI	Neutron star γ sources
< 92	95	⁵ AAD		ATLS	$pp \rightarrow jG$
< 72	95	⁶ CHATRCHYAN	11 U	CMS	$pp \rightarrow jG$
< 245	95	⁷ AALTONEN	O8AC	CDF	$p\overline{p} \rightarrow \gamma G, jG$
< 615	95	⁸ ABAZOV	085	D0	$p\overline{p} \rightarrow \gamma G$
< 0.916	95	⁹ DAS	80		Supernova cooling
< 350	95	¹⁰ ABULENCIA,A	06	CDF	$p\overline{p} \rightarrow jG$
< 270	95	¹¹ ABDALLAH	05 B	DLPH	$e^+e^- o \gamma G$
< 210	95	¹² ACHARD	04E	L3	$e^+e^- o \gamma G$
< 480	95	¹³ ACOSTA	04 C	CDF	$\overline{p}p \rightarrow jG$
< 0.00038	95	14 CASSE	04		Neutron star γ sources
< 610	95	15 ABAZOV	03	D0	$\overline{p}p \rightarrow jG$
< 0.96	95	¹⁶ HANNESTAD	03		Supernova cooling
< 0.096	95	¹⁷ HANNESTAD	03		Diffuse γ background
< 0.051	95	¹⁸ HANNESTAD	03		Neutron star γ sources
< 0.00016	95	¹⁹ HANNESTAD	03		Neutron star heating
< 300	95	²⁰ HEISTER	03 C	ALEP	$e^+e^- o \gamma G$
		²¹ FAIRBAIRN	01		Cosmology
< 0.66	95	²² HANHART	01		Supernova cooling
		²³ CASSISI	00		Red giants
<1300	95	²⁴ ACCIARRI	99 S	L3	$e^+e^- \rightarrow ZG$
1				. 1 .	

NODE=S071OS6

OCCUR=2 OCCUR=3 OCCUR=4

 1 CHATRCHYAN 12AP search for $pp \rightarrow jG$, using 5.0 fb $^{-1}$ of data at $\sqrt{s}=7$ TeV to place bounds on M_D for two to six extra dimensions, from which this bound on R is derived. See their Table 7 for bounds on all $\delta \leq 6$. AAD 13C search for $pp \to \gamma G$, using 4.6 fb⁻¹ of data at $\sqrt{s}=7$ TeV to place bounds

on M_D for two to six extra dimensions, from which this bound on R is derived.

 3 AAD $^-$ 13D search for the dijet decay of quantum black holes in 4.8 fb $^{-1}$ of data produced in pp collisions at $\sqrt{s}=7$ TeV to place bounds on M_D for two to seven extra dimensions, from which these bounds on R are derived. Limits on M_D for all $\delta \leq 7$ are given in

their Table 3. 4 AJELLO 12 obtain a limit on R from the gamma-ray emission of point γ sources that arise from the photon decay of KK gravitons which are gravitationally bound around neutron stars. Limits for all $\delta \leq 7$ are given in their Table 7.

⁵ AAD 11S search for $pp \to jG$, using 33 pb⁻¹ of data at $\sqrt{s} = 7$ TeV, to place bounds on M_D for two to four extra dimensions, from which these bounds on R are derived. See their Table 3 for bounds on all $\delta \le 4$.

NODE=S071OS6;LINKAGE=CC

NODE=S071OS6;LINKAGE=GA

NODE=S071OS6;LINKAGE=TA

NODE=S071OS6;LINKAGE=AJ

NODE=S071OS6;LINKAGE=DD

 6 CHATRCHYAN 11U search for $pp\to j\,G$, using 36 pb $^{-1}$ of data at $\sqrt{s}=7$ TeV, to place bounds on M_D for two to six extra dimensions, from which these bounds on R are derived. See their Table 3 for bounds on all $\delta~\leq~6$.

 7 AALTONEN 08AC search for $p\overline{p}\to\gamma\,G$ and $p\overline{p}\to j\,G$ at $\sqrt{s}=1.96$ TeV with 2.0 fb $^{-1}$ and 1.1 fb $^{-1}$ respectively, in order to place bounds on the fundamental scale and size of the extra dimensions. See their Table III for limits on all $\delta\leq 6$.

⁸ ABAZOV 08S search for $p\overline{p} \to \gamma G$, using 1 fb⁻¹ of data at $\sqrt{s}=1.96$ TeV to place bounds on M_D for two to eight extra dimensions, from which these bounds on R are derived. See their paper for intermediate values of δ .

 $^9\,\mathrm{DAS}$ 08 obtain a limit on R from Kaluza-Klein graviton cooling of SN1987A due to plasmon-plasmon annihilation.

 10 ABULENCIA,A 06 search for $p\overline{p}\to j\,G$ using 368 pb $^{-1}$ of data at $\sqrt{s}=1.96$ TeV. See their Table II for bounds for all $\delta\le 6$. 11 ABDALLAH 05B search for $e^+e^-\to \gamma\,G$ at $\sqrt{s}=180$ –209 GeV to place bounds on

¹¹ABDALLAH 05B search for $e^+e^- \to \gamma G$ at $\sqrt{s}=180$ –209 GeV to place bounds on the size of extra dimensions and the fundamental scale. Limits for all $\delta \leq 6$ are given in their Table 6. These limits supersede those in ABREU 00Z.

 12 ACHARD 04E search for $e^+\,e^-\to \gamma\, G$ at $\sqrt{s}=189$ –209 GeV to place bounds on the size of extra dimensions and the fundamental scale. See their Table 8 for limits with $\delta\,\leq\,$ 8. These limits supersede those in ACCIARRI 99R.

 $^-$ ACOSTA 04C search for $\overline{p}p \to jG$ at $\sqrt{s}=1.8$ TeV to place bounds on the size of extra dimensions and the fundamental scale. See their paper for bounds on $\delta=4,\,6.$

¹⁴CASSE 04 obtain a limit on R from the gamma-ray emission of point γ sources that arises from the photon decay of gravitons around newly born neutron stars, applying the technique of HANNESTAD 03 to neutron stars in the galactic bulge. Limits for all $\delta \leq 7$ are given in their Table I.

 15 ABAZOV 03 search for $p\overline{p}\to j\,G$ at $\sqrt{s}{=}1.8$ TeV to place bounds on M_D for 2 to 7 extra dimensions, from which these bounds on R are derived. See their paper for bounds on intermediate values of $\delta.$ We quote results without the approximate NLO scaling introduced in the paper.

 16 HANNESTAD 03 obtain a limit on R from graviton cooling of supernova SN1987a. Limits for all $\delta \leq 7$ are given in their Tables V and VI.

 17 HANNESTAD 03 obtain a limit on R from gravitons emitted in supernovae and which subsequently decay, contaminating the diffuse cosmic γ background. Limits for all $\delta \leq 7$ are given in their Tables V and VI. These limits supersede those in HANNESTAD 02.

 18 HANNESTAD 03 obtain a limit on R from gravitons emitted in two recent supernovae and which subsequently decay, creating point γ sources. Limits for all $\delta \leq 7$ are given in their Tables V and VI. These limits are corrected in the published erratum.

 19 HANNESTAD 03 obtain a limit on R from the heating of old neutron stars by the surrounding cloud of trapped KK gravitons. Limits for all $\delta \leq 7$ are given in their Tables V and VI. These limits supersede those in HANNESTAD 02.

²⁰ HEISTER 03C use the process $e^+e^- \to \gamma G$ at $\sqrt{s}=189$ –209 GeV to place bounds on the size of extra dimensions and the scale of gravity. See their Table 4 for limits with $\delta \leq 6$ for derived limits on M_D .

²¹ FAIRBAIRN 01 obtains bounds on R from over production of KK gravitons in the early universe. Bounds are quoted in paper in terms of fundamental scale of gravity. Bounds depend strongly on temperature of QCD phase transition and range from $R < 0.13~\mu \text{m}$ to $0.001~\mu \text{m}$ for $\delta = 2$; bounds for $\delta = 3,4$ can be derived from Table 1 in the paper.

 22 HANHART 01 obtain bounds on R from limits on graviton cooling of supernova SN 1987a using numerical simulations of proto-neutron star neutrino emission.

²³ CASSISI 00 obtain rough bounds on M_D (and thus R) from red giant cooling for δ =2,3. See their paper for details.

24 ACCIARRI 99s search for $e^+\,e^-\to~Z\,G$ at $\sqrt{s}{=}189$ GeV. Limits on the gravity scale are found in their Table 2, for $\delta \leq 4$.

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NODE=S071OS;LINKAGE=LO

NODE=S071OS;LINKAGE=BA

NODE=S071OS;LINKAGE=DA

NODE=S071OS6;LINKAGE=LE

NODE=S071OS;LINKAGE=AB

NODE=S071OS;LINKAGE=AR

NODE=S071OS;LINKAGE=AC

NODE=S071OS;LINKAGE=CA

NODE=S071OS;LINKAGE=ZB

NODE=S071OS;LINKAGE=HA

NODE=S071OS;LINKAGE=HB

NODE=S071OS;LINKAGE=HC

NODE=S071OS;LINKAGE=HD

NODE=S071OS;LINKAGE=3H

NODE=S071OS;LINKAGE=F

NODE=S071OS;LINKAGE=HT

NODE=S071OS;LINKAGE=CS

NODE=S071OS;LINKAGE=S9

Mass Limits on M_{TT}

This section includes limits on the cut-off mass scale, M_{TT} , of dimension-8 operators from KK graviton exchange in models of large extra dimensions. Ambiguities in the UV-divergent summation are absorbed into the parameter λ , which is taken to be $\lambda=\pm 1$ in the following analyses. Bounds for $\lambda=-1$ are shown in parenthesis after the bound for $\lambda=+1$, if appropriate. Different papers use slightly different definitions of the mass scale. The definition used here is related to another popular convention by $M_{TT}^4=(2/\pi)~\Lambda_{T}^4$, as discussed in the above Review on "Extra Dimensions."

VALUE (TeV)		CL%	DOCUMENT ID		TECN	COMMENT
> 3.2		95	¹ AAD	13E	ATLS	$pp \rightarrow e^+e^-, \mu^+\mu^-, \gamma\gamma$
• • • We	do not use	the follow	ving data for averag	ges, fi	ts, limits	, etc. • • •
> 2.66	(>2.27)	95		12Y	ATLS	$pp \rightarrow \gamma \gamma$
			³ BAAK	12	RVUE	Electroweak
> 2.86		95	⁴ CHATRCHYAN	l 12J	CMS	$pp ightarrow e^+e^-$, $\mu^+\mu^-$

NODE=S071GEX

NODE=S071GEX

NODE=S071GEX

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<sup>5</sup> CHATRCHYAN 12R CMS
> 2.84
                (>2.41)
                             95
                                         <sup>6</sup> AARON
                                                                                 e^{\pm} p \rightarrow e^{\pm} X
> 0.90
                                                               11c H1
                (>0.92)
                             95
                                         <sup>7</sup> CHATRCHYAN 11A CMS
> 1.74
                (>1.71)
                             95
                                                                                 pp \rightarrow \gamma \gamma
                                         <sup>8</sup> ABAZOV
                                                                                 p\overline{p} 	o 	ext{dijet}, ang. distrib.
                             95
                                                               09AE D0
> 1.48
                                         <sup>9</sup> ABAZOV
                                                                09D D0
                                                                                 p\overline{p} \rightarrow e^+e^-, \gamma\gamma
> 1.45
                             95
                                        <sup>10</sup> SCHAEL
                                                                                 e^+e^- 
ightarrow e^+e^-
                (> 1.0)
                             95
                                                                07A ALEP
> 1.1
                                        <sup>11</sup> ABDALLAH
                                                               06C DLPH e^+e^- \rightarrow \ell^+\ell^-
> 0.898
                (> 0.998) 95
                                        <sup>12</sup> GERDES
                                                                                 p\overline{p} \rightarrow e^+e^-, \gamma\gamma
> 0.853
                (> 0.939)95
                                                               06
                                        <sup>13</sup> ABAZOV
                                                                                 p\overline{p} \rightarrow \mu^+\mu^-
> 0.96
                (> 0.93)
                                                                05V
                                                                     D0
                                        <sup>14</sup> CHEKANOV
                                                               04B ZEUS
                                                                                 e^{\pm} p \rightarrow e^{\pm} X
> 0.78
                (> 0.79)
                             95
                                        <sup>15</sup> ABBIENDI
                                                                                e^+e^- \rightarrow \gamma \gamma
> 0.805
               (>0.956)95
                                                               03D OPAL
                                                                                 e^+ e^- \rightarrow ZZ
> 0.7
                (> 0.7)
                             95
                                        <sup>16</sup> ACHARD
                                                               03D L3
                (> 0.78)
                                        <sup>17</sup> ADLOFF
                                                               03
                                                                      H1
                                                                                 e^{\pm} p \rightarrow e^{\pm} X
> 0.82
                             95
> 1.28
                (>1.25)
                             95
                                        <sup>18</sup> GIUDICE
                                                               03
                                                                      RVUE
                                        <sup>19</sup> GIUDICE
>20.6
                (>15.7)
                             95
                                                               03
                                                                      RVUE Dim-6 operators
                                                                                                                              OCCUR=2
                                        <sup>20</sup> HEISTER
                                                                                e^+e^- \rightarrow \gamma \gamma
> 0.80
                (> 0.85)
                                                               03C ALEP
                             95
                                        <sup>21</sup> ACHARD
                                                               02D L3
                                                                                 e^+e^- \rightarrow \gamma\gamma
> 0.84
                (> 0.99)
                             95
                                        <sup>22</sup> ABBOTT
                                                               01 D0
                                                                                 p\overline{p} \rightarrow e^+e^-, \gamma\gamma
> 1.2
                (>1.1)
                             95
                                        <sup>23</sup> ABBIENDI
                                                               00R OPAL e^+e^- \rightarrow \mu^+\mu^-
                (> 0.63)
> 0.60
                             95
                                        <sup>23</sup> ABBIENDI
                                                               OOR OPAL e^+e^- \rightarrow \tau^+\tau^-
                                                                                                                              OCCUR=3
> 0.63
                (> 0.50)
                             95
                                        <sup>23</sup> ABBIENDI
                                                                00R OPAL e^+e^- \to \mu^+\mu^-, \tau^+\tau^-
                                                                                                                              OCCUR=5
> 0.68
                (> 0.61)
                                        <sup>24</sup> ABREU
                                                                00A DLPH e^+e^- \rightarrow \gamma \gamma
                                        <sup>25</sup> ABREU
> 0.680
                (>0.542) 95
                                                                00S DLPH e^+e^- \rightarrow \mu^+
                                                                                                                              OCCUR=6
                                        <sup>26</sup> CHANG
                                                                00B RVUE Electroweak
> 15-28
                             99.7
                                        <sup>27</sup> CHEUNG
                                                                      RVUE e^+e^- \rightarrow \gamma \gamma
> 0.98
                             95
                                                               00
                                        <sup>28</sup> GRAESSER
                                                                      RVUE (g-2)_{\mu}
> 0.29-0.38
                             95
                                                               00
                                        <sup>29</sup> HAN
                             95
                                                               00
                                                                      RVUE Electroweak
> 0.50-1.1
                                        <sup>30</sup> MATHEWS
                                                                      RVUE \overline{p}p \rightarrow jj
> 2.0
                (> 2.0)
                                                               00
                             95
                                                                      RVUE e^+e^- \rightarrow VV
                                        <sup>31</sup> MELE
                                                               00
> 1.0
                (> 1.1)
                             95
                                        <sup>32</sup> ABBIENDI
                                                                99P OPAL
                                        <sup>33</sup> ACCIARRI
                                                                99M L3
                                        <sup>34</sup> ACCIARRI
                                                                99s L3
                                        <sup>35</sup> BOURILKOV
               (>1.077) 95
                                                               99
> 1.412
   ^{1} AAD 13E use 4.9 and 5.0 fb^{-1} of data from \it pp collisions at \it \sqrt{s}= 7 TeV in the
```

 1 AAD 13E use 4.9 and 5.0 fb $^{-1}$ of data from pp collisions at $\sqrt{s}=7$ TeV in the dielectron and dimuon channels, respectively, to place lower limits on M_{TT} (equivalent to their $M_{\mbox{\scriptsize S}}).$ The dielectron and dimuon channels are combined with previous results in the diphoton channel to set the best limit. Bounds on individual channels and different priors can be found in their Table VIII.

 2 AAD 12Y use 2.12 fb $^{-1}$ of data from $p\,p$ collisions at $\sqrt{s}=7$ TeV to place lower limits on M_{TT} .

³ BAAK 12 use electroweak precision observables to place bounds on the ratio Λ_T/M_D as a function of M_D . See their Fig. 22 for constraints with a Higgs mass of 120 GeV.

 4 CHATRCHYAN 12J use approximately 2 fb $^{-1}$ of data from $p\,p$ collisions at $\sqrt{s}=7$ TeV in the dielectron and dimuon channels to place lower limits on Λ_{T} , here converted to M_{TT} .

 5 CHATRCHYAN 12R use 2.2 fb $^{-1}$ of data from pp collisions at $\sqrt{s}=7$ TeV to place lower limits on M_{TT} (equivalent to their $M_{\rm S}$).

⁶ AARON 11C search for deviations in the differential cross section of $e^{\pm}p \rightarrow e^{\pm}X$ in 446 pb⁻¹ of data taken at $\sqrt{s}=301$ and 319 GeV to place a bound on M_{TT} .

 7 CHATRCHYAN 11A use 36 pb $^{-1}$ of data from pp collisions at $\sqrt{s}=7$ TeV to place lower limits on Λ_T , here converted to M_{TT} .

 8 ABAZOV 09AE use dijet angular distributions in 0.7 fb $^{-1}$ of data from $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to place lower bounds on Λ_T (equivalent to their \textit{M}_S), here converted to \textit{M}_{TT} .

 9 ABAZOV 09D use 1.05 fb $^{-1}$ of data from $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to place lower bounds on Λ_T (equivalent to their M_s), here converted to M_{TT} .

 10 SCHAEL 07A use e^+e^- collisions at $\sqrt{s}=$ 189–209 GeV to place lower limits on \varLambda_T , here converted to limits on M_{TT} .

 11 ABDALLAH 06C use e^+e^- collisions at $\sqrt{s}\sim 130$ –207 GeV to place lower limits on M_{TT} , which is equivalent to their definition of $M_{\rm S}.$ Bound shown includes all possible final state leptons, $\ell=e,\,\mu,\,\tau.$ Bounds on individual leptonic final states can be found in their Table 31.

 12 GERDES 06 use 100 to 110 pb $^{-1}$ of data from $p\overline{p}$ collisions at $\sqrt{s}=1.8$ TeV, as recorded by the CDF Collaboration during Run I of the Tevatron. Bound shown includes a K-factor of 1.3. Bounds on individual e^+e^- and $\gamma\gamma$ final states are found in their Table I.

¹³ ABAZOV 05V use 246 pb⁻¹ of data from $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to search for deviations in the differential cross section to $\mu^+\mu^-$ from graviton exchange.

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NODE=S071GEX;LINKAGE=GT

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NODE=S071GEX;LINKAGE=CT

NODE=S071GEX;LINKAGE=YA

NODE=S071GEX;LINKAGE=AA

NODE=S071GEX;LINKAGE=CA

NODE=S071GEX;LINKAGE=ZO

NODE=S071GEX;LINKAGE=BA

NODE=S071GEX;LINKAGE=SC

NODE=S071GEX;LINKAGE=BD

NODE=S071GEX;LINKAGE=GE

NODE=S071GEX;LINKAGE=AZ

- ¹⁴CHEKANOV 04B search for deviations in the differential cross section of $e^{\pm}p \rightarrow e^{\pm}X$ with 130 pb^{-1} of combined data and Q^2 values up to 40,000 GeV² to place a bound on M_{TT} .
- 15 ABBIENDI 03D use $e^+\,e^-$ collisions at $\sqrt{s}{=}181{-}209$ GeV to place bounds on the ultraviolet scale M_{TT} , which is equivalent to their definition of M_{S} .
- ¹⁶ ACHARD 03D look for deviations in the cross section for $e^+e^- \to ZZ$ from $\sqrt{s}=200$ –209 GeV to place a bound on M_{TT} .
- ¹⁷ ADLOFF 03 search for deviations in the differential cross section of $e^{\pm}p \rightarrow e^{\pm}X$ at \sqrt{s} =301 and 319 GeV to place bounds on M_{TT} .
- $^{18}\,\mathrm{GIUDICE}$ 03 review existing experimental bounds on M_{TT} and derive a combined limit.
- 19 GIUDICE 03 place bounds on Λ_6 , the coefficient of the gravitationally-induced dimension- 6 operator $(2\pi\lambda/\Lambda_6^2)(\sum\overline{f}\gamma_\mu\gamma^5f)(\sum\overline{f}\gamma^\mu\gamma^5f)$, using data from a variety of experiments. Results are quoted for $\lambda=\pm1$ and are independent of δ .
- $^{20}\, {\rm HEISTER}$ 03C use $e^+\,e^-$ collisions at $\sqrt{s} = 189 209$ GeV to place bounds on the scale of dim-8 gravitational interactions. Their M_S^\pm is equivalent to our M_{TT} with $\lambda = \pm 1.$
- 21 ACHARD 02 search for s-channel graviton exchange effects in $\rm e^+\,e^-\to\gamma\gamma$ at $\rm E_{cm}=192-209~GeV.$
- ²² ABBOTT 01 search for variations in differential cross sections to e⁺ e⁻ and $\gamma\gamma$ final states at the Tevatron.
- ²³ ABBIENDI 00R uses e^+e^- collisions at $\sqrt{s}=$ 189 GeV.
- ²⁴ ABREU 00A search for s-channel graviton exchange effects in e⁺ e⁻ $\to \gamma\gamma$ at $E_{\rm cm}=$ 189–202 GeV.
- 25 ABREU 00s uses $e^+\,e^-$ collisions at $\sqrt{s}{=}183$ and 189 GeV. Bounds on μ and τ individual final states given in paper.
- ²⁶CHANG 00B derive 3σ limit on M_{TT} of (28,19,15) TeV for δ =(2,4,6) respectively assuming the presence of a torsional coupling in the gravitational action. Highly model dependent.
- 27 CHEUNG 00 obtains limits from anomalous diphoton production at OPAL due to graviton exchange. Original limit for $\delta{=}4.$ However, unknown UV theory renders δ dependence unreliable. Original paper works in HLZ convention.
- 28 GRAESSER 00 obtains a bound from graviton contributions to g-2 of the muon through loops of 0.29 TeV for $\delta{=}2$ and 0.38 TeV for $\delta{=}4,6$. Limits scale as $\lambda^{1/2}$. However calculational scheme not well-defined without specification of high-scale theory. See the "Extra Dimensions Review."
- HAN 00 calculates corrections to gauge boson self-energies from KK graviton loops and constrain them using S and T. Bounds on M_{TT} range from 0.5 TeV (δ =6) to 1.1 TeV (δ =2); see text. Limits have strong dependence, $\lambda^{\delta+2}$, on unknown λ coefficient.
- 30 MATHEWS 00 search for evidence of graviton exchange in CDF and DØ dijet production data. See their Table 2 for slightly stronger δ -dependent bounds. Limits expressed in terms of $\widetilde{M}_5^4 = M_{TT}^4/8$.
- ³¹ MELE 00 obtains bound from KK graviton contributions to $e^+e^- \rightarrow VV$ ($V=\gamma,W,Z$) at LEP. Authors use Hewett conventions.
- at LEP. Authors use riewett conventions. 32 ABBIENDI 99P search for s-channel graviton exchange effects in $e^+\,e^- \to \gamma\gamma$ at $E_{\rm cm}=$ 189 GeV. The limits $G_+ >$ 660 GeV and $G_- >$ 634 GeV are obtained from combined $E_{\rm cm}=$ 183 and 189 GeV data, where G_\pm is a scale related to the fundamental gravity scale.
- ³³ ACCIARRI 99M search for the reaction $e^+e^- \to \gamma G$ and s-channel graviton exchange effects in $e^+e^- \to \gamma \gamma$, W^+W^- , ZZ, e^+e^- , $\mu^+\mu^-$, $\tau^+\tau^-$, $q\overline{q}$ at $E_{\rm cm}=183$ GeV. Limits on the gravity scale are listed in their Tables 1 and 2.
- ³⁴ ACCIARRI 99S search for the reaction $e^+e^- \to ZG$ and s-channel graviton exchange effects in $e^+e^- \to \gamma\gamma$, W^+W^- , ZZ, e^+e^- , $\mu^+\mu^-$, $\tau^+\tau^-$, $q\overline{q}$ at $E_{\rm cm}=$ 189 GeV. Limits on the gravity scale are listed in their Tables 1 and 2.
- 35 BOURILKOV 99 performs global analysis of LEP data on e $^+\,e^-$ collisions at $\sqrt{s}{=}183$ and 189 GeV. Bound is on Λ_T .

Limits on $1/R = M_c$

This section includes limits on $1/R=M_{\rm C}$, the compactification scale in models with TeV extra dimensions, due to exchange of Standard Model KK excitations. Bounds assume fermions are not in the bulk, unless stated otherwise. See the "Extra Dimensions" review for discussion of model dependence.

VALUE (TeV)CL%DOCUMENT IDTECNCOMMENT>4.1695 1 AAD12CC ATLS $pp \rightarrow \ell \bar{\ell}$

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NODE=S071GEX;LINKAGE=AP

NODE=S071GEX;LINKAGE=B3

NODE=S071GEX;LINKAGE=DQ

 ${\sf NODE}{=}{\sf S071GEX;LINKAGE}{=}{\sf C}$

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>1.40	95	² AAD	12CP ATLS	$pp \rightarrow \gamma \gamma, \delta = 6, M_D = 5 \text{ TeV}$
>1.23	95	³ AAD		$pp \rightarrow \gamma \gamma$, δ =6, M_D =5 TeV
>0.26	95	⁴ ABAZOV	12M D0	$p\overline{p} \rightarrow \mu\mu$
>0.75	95	⁵ BAAK	12 RVUE	Electroweak
		⁶ FLACKE	12 RVUE	Electroweak
>0.43	95	⁷ NISHIWAKI	12 RVUE	$H ightarrow \ W W$, $\gamma \gamma$
>0.729	95	⁸ AAD	11F ATLS	$pp \rightarrow \gamma\gamma$, δ =6, M_D =5 TeV
>0.961	95	⁹ AAD	11X ATLS	$pp \rightarrow \gamma \gamma$, δ =6, M_D =5 TeV
>0.477	95	¹⁰ ABAZOV	10P D0	$p\overline{p} \rightarrow \gamma\gamma$, $\delta=6$, $M_D=5$ TeV
>1.59	95	¹¹ ABAZOV	09AE D0	$p\overline{p} \rightarrow \text{dijet}$, angular dist.
>0.6	95	¹² HAISCH	07 RVUE	$\overline{B} \rightarrow X_{s} \gamma$
>0.6	90	¹³ GOGOLADZE	06 RVUE	Electroweak
>3.3	95	¹⁴ CORNET	00 RVUE	Electroweak
> 3.3–3.8	95	¹⁵ RIZZO	00 RVUE	Electroweak

 1 AAD 12CC use 4.9 and 5.0 fb $^{-1}$ of data from pp collisions at $\sqrt{s}=7$ TeV in the dielectron and dimuon channels, respectively, to place a lower bound on the mass of the lightest KK Z/γ boson (equivalent to $1/R=M_{\rm C}$). The limit quoted here assumes a flat prior corresponding to when the pure Z/γ KK cross section term dominates. See their Section 15 for more details.

 2 AAD 12CP use diphoton events with large missing transverse momentum in 4.8 fb $^{-1}$ of data produced from pp collisions at $\sqrt{s}=7$ TeV to place a lower bound on the compactification scale in a universal extra dimension model with gravitational decays. The bound assumes that the cutoff scale \varLambda , for the radiative corrections to the Kaluza-Klein masses, satisfies $\Lambda/M_C=20$. The model parameters are chosen such that the decay $\gamma^*\to G\gamma$ occurs with an appreciable branching fraction.

 3 AAD $^{12}{\rm X}$ use diphoton events with large missing transverse momentum in $1.07~{\rm fb}^{-1}$ of data produced from pp collisions at $\sqrt{s}=7~{\rm TeV}$ to place a lower bound on the compactification scale in a universal extra dimension model with gravitational decays. The bound assumes that the cutoff scale \varLambda , for the radiative corrections to the Kaluza-Klein masses, satisfies $\varLambda/M_C=20$. The model parameters are chosen such that the decay $\gamma^*\to G\gamma$ occurs with an appreciable branching fraction.

⁴ ABAZOV 12M use same-sign dimuon events in 7.3 fb⁻¹ of data from $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to place a lower bound on the compactification scale 1/R, in models with universal extra dimensions where all Standard Model fields propagate in the bulk.

 $^5\,\mathrm{BAAK}$ 12 use electroweak precision observables to place a lower bound on the compactification scale 1/R, in models with universal extra dimensions and Standard Model fields propagating in the bulk. Bound assumes a 125 GeV Higgs mass. See their Fig. 25 for the bound as a function of the Higgs mass.

 6 FLACKE 12 use electroweak precision observables to place a lower bound on the compactification scale 1/R, in models with universal extra dimensions and Standard Model fields propagating in the bulk. See their Fig. 1 for the bound as a function of the universal bulk fermion mass parameter $\mu.$

 7 NISHIWAKI 12 use up to 2 fb $^{-1}$ of data from the ATLAS and CMS experiments that constrains the production cross section of a Higgs-like particle to place a lower bound on the compactification scale 1/R in universal extra dimension models. The quoted bound assumes Standard Model fields propagating in the bulk and a 125 GeV Higgs mass. See their Fig. 1 for the bound as a function of the Higgs mass.

 8 AAD 11F use diphoton events with large missing transverse energy in 3.1 pb $^{-1}$ of data produced from pp collisions at $\sqrt{s}=7$ TeV to place a lower bound on the compactification scale in a universal extra dimension model with gravitational decays. The bound assumes that the cutoff scale \varLambda , for the radiative corrections to the Kaluza-Klein masses, satisfies $\varLambda/M_c=20$. The model parameters are chosen such that the decay $\gamma^*\to G\gamma$ occurs with an appreciable branching fraction.

 9 AAD 11X use diphoton events with large missing transverse energy in 36 pb $^{-1}$ of data produced from pp collisions at $\sqrt{s}=7$ TeV to place a lower bound on the compactification scale in a universal extra dimension model with gravitational decays. The bound assumes that the cutoff scale Λ , for the radiative corrections to the Kaluza-Klein masses, satisfies $\Lambda/M_C=20$. The model parameters are chosen such that the decay $\gamma^*\to G\gamma$ occurs with an appreciable branching fraction.

 10 ABAZOV 10P use diphoton events with large missing transverse energy in 6.3 fb $^{-1}$ of data produced from $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to place a lower bound on the compactification scale in a universal extra dimension model with gravitational decays. The bound assumes that the cutoff scale $\Lambda,$ for the radiative corrections to the Kaluza-Klein masses, satisfies $\Lambda/\mathrm{M}_c{=}20.$ The model parameters are chosen such that the decay

 $\gamma^* \to G\gamma$ occurs with an appreciable branching fraction.

 11 ABAZOV 09AE use dijet angular distributions in 0.7 fb $^{-1}$ of data from $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to place a lower bound on the compactification scale.

 $^{12}\,\mbox{HAISCH}$ 07 use inclusive $\overline{B}\mbox{-meson}$ decays to place a Higgs mass independent bound on the compactification scale 1/R in the minimal universal extra dimension model.

13 GOGOLADZE 06 use electroweak precision observables to place a lower bound on the compactification scale in models with universal extra dimensions. Bound assumes a 115 GeV Higgs mass. See their Fig. 3 for the bound as a function of the Higgs mass.

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NODE=S071KK;LINKAGE=ZO

NODE=S071KK;LINKAGE=HA

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 14 CORNET 00 translates a bound on the coefficient of the 4-fermion operator $(\bar\ell\gamma_\mu\tau^a\ell)(\bar\ell\gamma^\mu\tau^a\ell)$ derived by Hagiwara and Matsumoto into a limit on the mass scale ... of KK W bosons.

15 RIZZO 00 obtains limits from global electroweak fits in models with a Higgs in the bulk (3.8 TeV) or on the standard brane (3.3 TeV).

NODE=S071KK;LINKAGE=B

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Limits on Kaluza-Klein Gravitons in Warped Extra Dimensions

This sections places limits on the mass of the first Kaluza-Klein (KK) excitation of the graviton in the warped extra dimension model of Randall and Sundrum. Bounds in parenthesis assume Standard Model fields propagate in the bulk. Experimental bounds depend strongly on the warp parameter, k. See the "Extra Dimensions" review for a full discussion.

Here we list limits for the value of the warp parameter $k/\overline{M}_P=0.1$.

> 2.160 × 10³ (CL = 95%) [>1.630 × 10³ GeV (CL = 95%) OUR 2012 BEST LIMIT] >2160 95 1 AAD 12CC ATLS $pp \rightarrow G \rightarrow \ell \bar{\ell}$ • • • We do not use the following data for averages, fits, limits, etc. • • • >1230 (>840) 95 2 AAD 13A ATLS $pp \rightarrow G \rightarrow WW$ > 845 95 3 AAD 12AD ATLS $pp \rightarrow G \rightarrow ZZ$ >1950 95 4 AAD 12Y ATLS $pp \rightarrow G \rightarrow ZZ$ 6 BAAK 12 RVUE Electroweak >1840 95 7 CHATRCHYAN12R CMS $pp \rightarrow G \rightarrow ZZ$ 9 AALTONEN 11G CDF $p\bar{p} \rightarrow G \rightarrow ZZ$ >1058 95 4 AAD 11AD ATLS $pp \rightarrow G \rightarrow ZZ$ 6 BAAK 12 RVUE $pp \rightarrow G \rightarrow ZZ$ 11AD ATLS $pp \rightarrow G \rightarrow ZZ$ 12ALTONEN 11G CDF $p\bar{p} \rightarrow G \rightarrow ZZ$ >1058 95 10 AALTONEN 11G CDF $p\bar{p} \rightarrow G \rightarrow ZZ$ >1079 21 ABAZOV 11H D0 $p\bar{p} \rightarrow G \rightarrow \ell \bar{\ell}$ >607 13 AALTONEN 10N CDF $p\bar{p} \rightarrow G \rightarrow \ell \bar{\ell}$ 14 ABAZOV 11C DO 15 AALTONEN 10N CDF $p\bar{p} \rightarrow G \rightarrow \ell \bar{\ell}$ >900 16 ABAZOV 08J D0 $p\bar{p} \rightarrow G \rightarrow \ell \bar{\ell}$ 17 AALTONEN 07G CDF $p\bar{p} \rightarrow G \rightarrow \ell \bar{\ell}$ >889 18 AALTONEN 07H CDF $p\bar{p} \rightarrow G \rightarrow \ell \bar{\ell}$ 19 ABAZOV 05N D0 $p\bar{p} \rightarrow G \rightarrow \ell \bar{\ell}$ >889 785 19 ABAZOV 05N D0 $p\bar{p} \rightarrow G \rightarrow \ell \bar{\ell}$ 9 ABULENCIA 05A CDF $p\bar{p} \rightarrow G \rightarrow \ell \bar{\ell}$	VALUE (GeV)		CL%	DOCUMENT ID		TECN	COMMENT
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$\begin{array}{cccccccccccccccccccccccccccccccccccc$	> 607				10N (CDF	$p\overline{p} \rightarrow G \rightarrow WW$
> 900 $ \begin{array}{ccccccccccccccccccccccccccccccccccc$	>1050				10F [D0	$p\overline{p} \rightarrow G \rightarrow e^+e^-, \gamma\gamma$
$\begin{array}{cccccccccccccccccccccccccccccccccccc$					08s (CDF	$p\overline{p} \rightarrow G \rightarrow ZZ$
$>$ 889 $>$ 18 AALTONEN 07H CDF $p\overline{p} \rightarrow G \rightarrow e\overline{e}$ $>$ 785 19 ABAZOV 05N D0 $p\overline{p} \rightarrow G \rightarrow \ell\ell, \gamma\gamma$	> 900				1 L80	D0	$p\overline{p} \rightarrow G \rightarrow e^+e^-, \gamma\gamma$
$>$ 785 19 ABAZOV 05N D0 $p\overline{p} \rightarrow G \rightarrow \ell\ell, \gamma\gamma$					07G (CDF	$p\overline{p} \rightarrow G \rightarrow \gamma\gamma$
	> 889				07н (CDF	$p\overline{p} \rightarrow G \rightarrow e\overline{e}$
>710 ABULENCIA 05A CDF $p\overline{p} ightarrow G ightarrow \ell \overline{\ell}$	> 785				05N [D0	$p\overline{p} ightarrow G ightarrow \ell\ell$, $\gamma\gamma$
	> 710			²⁰ ABULENCIA	05A (CDF	$p\overline{p} ightarrow G ightarrow \ell \overline{\ell}$

 1 AAD 12CC use 4.9 and 5.0 fb $^{-1}$ of data from pp collisions at $\sqrt{s}=7$ TeV in the dielectron and dimuon channels, respectively, to place a lower bound on the mass of the lightest KK graviton. See their Figure 5 for limits on the lightest KK graviton mass as a function of k/\overline{M}_P .

 2 AAD 13A use 4.7 fb $^{-1}$ of data from $p\,p$ collisions at $\sqrt{s}=7$ TeV to place a lower bound on the mass of the lightest KK graviton.

³ AAD 12AD use $1.02~{\rm fb}^{-1}$ of data from pp collisions at $\sqrt{s}=7~{\rm TeV}$ to search for KK gravitons in a warped extra dimension decaying to ZZ dibosons in the IIjj and IIII channels ($\ell = e, \mu$). The limit is quoted for the combined IIjj + IIII channels. See their Figure 5 for limits on the cross section $\sigma(G \to ZZ)$ as a function of the graviton mass.

 4 AAD 12Y use 2.12 fb $^{-1}$ of data from $p\,p$ collisions at $\sqrt{s}=7$ TeV in the diphoton channel to place lower limits on the mass of the lightest KK graviton. The diphoton channel is combined with previous results in the dielectric and dimuon channels to set the best limit. See their Table 3 for warp parameter values k/\overline{M}_P between 0.01 and 0.1.

⁵ AALTONEN 12V use 6 fb⁻¹ of data from $p\bar{p}$ collisions at $\sqrt{s}=1.96$ TeV to search for KK gravitons in a warped extra dimension decaying to ZZ dibosons in the IIjj and IIII channels ($\ell=e, \mu$). It provides improved limits over the previous analysis in AALTONEN 11G. See their Figure 16 for limits from all channels combined on the cross section times branching ratio $\sigma(p\bar{p}\to G^*\to ZZ)$ as a function of the graviton mass.

 6 BAAK 12 use electroweak precision observables to place a lower bound on the compactification scale $k~e^{-\pi~k~R}$, assuming Standard Model fields propagate in the bulk and the Higgs is confined to the IR brane. See their Fig. 27 for more details. 7 CHATRCHYAN 12R use 2.2 fb $^{-1}$ of data from pp collisions at $\sqrt{s}=7$ TeV in the

 7 CHATRCHYAN 12R use 2.2 fb $^{-1}$ of data from pp collisions at $\sqrt{s}=7$ TeV in the diphoton channel to place lower limits on the mass of the lightest KK graviton. See their Table III for warp parameter values k/\overline{M}_P between 0.01 and 0.1.

⁸ AAD 11AD use 1.08 and 1.21 fb⁻¹ of data from pp collisions at $\sqrt{s}=7$ TeV in the dielectron and dimuon channels, respectively, to place a lower bound on the mass of the lightest graviton. For warp parameter values k/\overline{M}_P between 0.01 to 0.1 the lower limit on the mass of the lightest graviton is between 0.71 and 1.63 TeV. See their Table IV for more datails

 9 AALTONEN 11G use 2.5–2.9 fb $^{-1}$ of data from $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to search for KK gravitons in a warped extra dimension decaying to ZZ dibosons via the

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NODE=S071RSG;LINKAGE=YA

NODE=S071RSG;LINKAGE=DA

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eeee, ee $\mu\mu$, $\mu\mu\mu\mu$, eejj, and $\mu\mu$ jj channels. See their Fig. 20 for limits on the cross section $\sigma(G\to ZZ)$ as a function of the graviton mass.

 10 AALTONEN 11R uses $5.7~{\rm fb}^{-1}$ of data from $p\overline{p}$ collisions at $\sqrt{s}=1.96~{\rm TeV}$ in the dielectron channel to place a lower bound on the mass of the lightest graviton. It provides combined limits with the diphoton channel analysis of AALTONEN 11U. For warp parameter values k/\overline{M}_P between 0.01 to 0.1 the lower limit on the mass of the lightest graviton is between 612 and 1058 GeV. See their Table I for more details.

 11 ABAZOV 11H use 5.4 fb $^{-1}$ of data from $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to place a lower bound on the mass of the lightest graviton. Their 95% C.L. exclusion limit does not include masses less than 300 GeV.

12 CHATRCHYAN 11 use 35 and 40 pb⁻¹ of data from pp collisions at $\sqrt{s}=7$ TeV in the dielectron and dimuon channels, respectively, to place a lower bound on the mass of the lightest graviton. For a warp parameter value $k/\overline{M}_P=0.05$, the lower limit on the mass of the lightest graviton is 0.855 TeV.

 13 AALTONEN 10N use 2.9 fb $^{-1}$ of data from $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to place a lower bound on the mass of the lightest graviton.

 14 ABAZOV 10F use 5.4 fb $^{-1}$ of data from $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to place a lower bound on the mass of the lightest graviton. For warp parameter values of k/\overline{M}_P between 0.01 and 0.1 the lower limit on the mass of the lightest graviton is between 560 and 1050 GeV. See their Fig. 3 for more details.

 15 AALTONEN 08s use $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to search for KK gravitons in warped extra dimensions. They search for graviton resonances decaying to four electrons via two Z bosons using 1.1 fb $^{-1}$ of data. See their Fig. 8 for limits on $\sigma \cdot \mathrm{B}(G \to ZZ)$ versus the graviton mass.

 16 ABAZOV 08 J use $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to search for KK gravitons in warped extra dimensions. They search for graviton resonances decaying to electrons and photons using $1~{\rm fb}^{-1}$ of data. For warp parameter values of k/\overline{M}_P between 0.01 and 0.1 the lower limit on the mass of the lightest excitation is between 300 and 900 GeV. See their Fig. 4 for more details.

¹⁷ AALTONEN 07G use $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to search for KK gravitons in warped extra dimensions. They search for graviton resonances decaying to photons using 1.2 fb $^{-1}$ of data. For warp parameter values of $k/\overline{M}_P=0.1$, 0.05, and 0.01 the bounds on the graviton mass are 850, 694, and 230 GeV, respectively. See their Fig. 3 for more details. See also AALTONEN 07H.

18 AALTONEN 07H use $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to search for KK gravitons in warped extra dimensions. They search for graviton resonances decaying to electrons using $1.3~{\rm fb^{-1}}$ of data. For a warp parameter value of $k/\overline{M}_P=0.1$ the bound on the graviton mass is 807 GeV. See their Fig. 4 for more details. A combined analysis with the diphoton data of AALTONEN 07G yields for $k/\overline{M}_P=0.1$ a graviton mass lower bound of 889 GeV.

 19 ABAZOV 05N use $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to search for KK gravitons in warped extra dimensions. They search for graviton resonances decaying to muons, electrons or photons, using 260 pb $^{-1}$ of data. For warp parameter values of $k/\overline{M}_P=0.1$, 0.05, and 0.01, the bounds on the graviton mass are 785, 650 and 250 GeV respectively. See their Fig. 3 for more details.

²⁰ ABULENCIA 05A use $p\overline{p}$ collisions at $\sqrt{s}=1.96$ TeV to search for KK gravitons in warped extra dimensions. They search for graviton resonances decaying to muons or electrons, using 200 pb⁻¹ of data. For warp parameter values of $k/\overline{M}p=0.1, 0.05$, and 0.01, the bounds on the graviton mass are 710, 510 and 170 GeV respectively.

NODE=S071RSG;LINKAGE=AT

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NODE=S071RSG;LINKAGE=AA

NODE=S071RSG;LINKAGE=AL

NODE=S071RSG;LINKAGE=AB

NODE=S071RSG;LINKAGE=AU

REFERENCES FOR Extra Dimensions

AAD	13A	PL B718 860	G. Aad et al.	(ATLAS Collab.)	
AAD	13C	PRL 110 011802	G. Aad et al.	(ATLAS Collab.)	
AAD	13D	JHEP 1301 029	G. Aad et al.	(ATLAS Collab.)	
AAD	13E	PR D87 015010	G. Aad et al.	(ATLAS Collab.)	
AAD	12AD	PL B712 331	G. Aad et al.	(ATLAS Collab.)	
AAD	12CC	JHEP 1211 138	G. Aad et al.	(ATLAS Collab.)	
AAD	12CP	PL B718 411	G. Aad et al.	(ATLAS Collab.)	
AAD	12X	PL B710 519	G. Aad et al.	(ATLAS Collab.)	
AAD	12Y	PL B710 538	G. Aad et al.	(ATLAS Collab.)	
AALTONEN	12V	PR D85 012008	T. Aaltonen et al.	` (CDF Collab.)	
ABAZOV	12M	PRL 108 131802	V.M. Abazov et al.	`(D0 Collab.)	
AJELLO	12	JCAP 1202 012	M. Ajello et al.	(Fermi-LAT Collab.)	
BAAK	12	EPJ C72 2003	M. Baak et al.	` (Gfitter Group)	
CHATRCHYAN	12AP	JHEP 1209 094	S. Chatrchyan et al.	(CMS Collab.)	
CHATRCHYAN	12J	PL B711 15	S. Chatrchyan et al.	(CMS Collab.)	
CHATRCHYAN	12R	PRL 108 111801	S. Chatrchyan et al.	(CMS Collab.)	
FLACKE	12	PR D85 126007	T. Flacke, C. Pasold	` (WURZ)	
NISHIWAKI	12	PL B707 506	K. Nishiwaki et al.	(KOBE, OSAK)	
AAD		PRL 107 272002	G. Aad <i>et al.</i>	(ATLAS Collab.)	
AAD	11F		G. Aad <i>et al.</i>	(ATLAS Collab.)	
AAD	11S	PL B705 294	G. Aad <i>et al.</i>	(ATLAS Collab.)	
AAD	11X	EPJ C71 1744	G. Aad <i>et al.</i>	(ATLAS Collab.)	
AALTONEN	11G	PR D83 112008	T. Aaltonen <i>et al.</i>	(CDF Collab.)	
AALTONEN	11R	PRL 107 051801	T. Aaltonen <i>et al.</i>	(CDF Collab.)	
AALTONEN	11U	PR D83 011102	T. Aaltonen <i>et al.</i>	(CDF Collab.)	
AARON	11C	PL B705 52	F. D. Aaron <i>et al.</i>	(H1 Collab.)	
ABAZOV	11H	PRL 107 011801	V. M. Abazov et al.	(D0 Collab.)	
BEZERRA	11	PR D83 075004	V.B. Bezerra et al.		
CHATRCHYAN		JHEP 1105 093	S. Chatrchyan et al.	(CMS Collab.)	
CHATRCHYAN		JHEP 1105 085	S. Chatrchyan et al.	(CMS Collab.)	
CHATRCHYAN		PRL 107 201804	S. Chatychyan et al.	(CMS Collab.)	
SUSHKOV	11	PRL 107 171101	A.O. Sushkov et al.		

NODE=S071 REFID=54789

						DEELD FOOTO
AALTONEN	10N	PRL 104 241801	T. Aaltonen <i>et al.</i>	(CDF	Collab.)	REFID=53313
ABAZOV	10F	PRL 104 241802	V.M. Abazov <i>et al.</i>	(D0	Collab.)	REFID=53318
ABAZOV	10P	PRL 105 221802	V.M. Abazov et al.	(D0	Collab.)	REFID=53443
BEZERRA	10	PR D81 055003	V.B. Bezerra et al.	(-	/	REFID=53391
ABAZOV		PRL 103 191803	V.M. Abazov et al.	(D0	Collab.)	REFID=53064
ABAZOV	09D	PRL 102 051601	V.M. Abazov et al.	(D0	Collab.)	REFID=52649
MASUDA	09	PRL 102 171101	M. Masuda, M. Sasaki		(ICRR)	REFID=52844
AALTONEN	08AC	PRL 101 181602	T. Aaltonen <i>et al.</i>	(CDF	Collab.)	REFID=52563
AALTONEN	08S	PR D78 012008	T. Aaltonen <i>et al.</i>	(CDF	Collab.)	REFID=52373
ABAZOV	08J	PRL 100 091802	V.M. Abazov et al.	`(D0	Collab.)	REFID=52388
ABAZOV	085	PRL 101 011601	V.M. Abazov et al.		Collab.)	REFID=52397
DAS	08	PR D78 063011			conab.)	REFID=52530
			P.K. Das, V.H.S. Kumar, P.K. Suresh		(CTANI)	
GERACI	80	PR D78 022002	A.A. Geraci <i>et al.</i>		(STAN)	REFID=52417
TRENKEL	80	PR D77 122001	C. Trenkel			REFID=52408
AALTONEN	07G	PRL 99 171801	T. Aaltonen <i>et al.</i>	(CDF	Collab.)	REFID=52012
AALTONEN	07H	PRL 99 171802	T. Aaltonen <i>et al.</i>	(CDF	Collab.)	REFID=52013
DECCA	07A	EPJ C51 963	R.S. Decca et al.	,	,	REFID=51880
HAISCH	07	PR D76 034014	U. Haisch, A. Weiler			REFID=51860
KAPNER	07	PRL 98 021101	D.J. Kapner <i>et al.</i>			REFID=51607
				(ALEDII	C.II.L.)	
SCHAEL	07A	EPJ C49 411	S. Schael et al.	(ALEPH	Collab.)	REFID=51743
TU	07	PRL 98 201101	LC. Tu <i>et al.</i>			REFID=51808
ABDALLAH	06C	EPJ C45 589	J. Abdallah <i>et al.</i>	(DELPHI	Collab.)	REFID=51224
ABULENCIA, A	06	PRL 97 171802	A. Abulencia et al.	(CDF	Collab.)	REFID=51434
GERDES	06	PR D73 112008	D. Gerdes et al.	`	,	REFID=51286
GOGOLADZE	06	PR D74 093012	I. Gogoladze, C. Macesanu			REFID=51544
				(D0	Callab)	
ABAZOV	05N	PRL 95 091801	V.M. Abazov et al.		Collab.)	REFID=50740
ABAZOV	05V	PRL 95 161602	V.M. Abazov et al.		Collab.)	REFID=50927
ABDALLAH	05B	EPJ C38 395	J. Abdallah <i>et al.</i>	(DELPHI	Collab.)	REFID=50448
ABULENCIA	05A	PRL 95 252001	A. Abulencia <i>et al.</i>	(CDF	Collab.)	REFID=51009
SMULLIN	05	PR D72 122001	S.J. Smullin et al.	,	,	REFID=51007
ACHARD	04E	PL B587 16	P. Achard et al.	(1.3	Collab.)	REFID=49893
ACOSTA	04C	PRL 92 121802	D. Acosta et al.		Collab.)	REFID=49935
				(CDI	Collab.)	
CASSE	04	PRL 92 111102	M. Casse et al.	(ZEUC	C !! !)	REFID=50055
CHEKANOV	04B	PL B591 23	S. Chekanov et al.		Collab.)	REFID=49912
HOYLE	04	PR D70 042004	C.D. Hoyle <i>et al.</i>		(WASH)	REFID=51084
ABAZOV	03	PRL 90 251802	V.M. Abazov <i>et al.</i>	(D0	Collab.)	REFID=49443
ABBIENDI	03D	EPJ C26 331	G. Abbiendi et al.	(OPAL	Collab.)	REFID=49290
ACHARD	03D	PL B572 133	P. Achard et al.	`	Collab.)	REFID=49556
ADLOFF	03	PL B568 35	C. Adloff <i>et al</i> .		Collab.)	REFID=49522
				(111	Collab.)	REFID=49357
CHIAVERINI	03	PRL 90 151101	J. Chiaverini et al.			
GIUDICE	03	NP B663 377	G.F. Giudice, A. Strumia			REFID=49427
HANNESTAD	03	PR D67 125008	S. Hannestad, G.G. Raffelt			REFID=49536
Also		PR D69 029901(errat)	S. Hannestad, G.G. Raffelt			REFID=50033
HEISTER	03C	EPJ C28 1 ` ´	A. Heister et al.	(ALEPH	Collab.)	REFID=49379
LONG	03	Nature 421 922	J.C. Long et al.	`	,	REFID=49699
ACHARD	02	PL B524 65	P. Achard <i>et al.</i>	(13	Collab.)	REFID=48524
ACHARD	02D	PL B531 28	P. Achard et al.	(LS	Collab.)	REFID=48663
HANNESTAD	02	PRL 88 071301	S. Hannestad, G. Raffelt			REFID=48698
ABBOTT	01	PRL 86 1156	B. Abbott <i>et al.</i>	(D0	Collab.)	REFID=48041
FAIRBAIRN	01	PL B508 335	M. Fairbairn			REFID=48264
HANHART	01	PL B509 1	C. Hanhart et al.			REFID=48160
HOYLE	01	PRL 86 1418	C.D. Hoyle et al.			REFID=48082
ABBIENDI	00R	EPJ C13 553	G. Abbiendi <i>et al.</i>	(OPAI	Collab.)	REFID=47644
ABREU	00A	PL B491 67	P. Abreu <i>et al.</i>	(DELPHI		REFID=47825
ABREU	00S	PL B485 45	P. Abreu et al.	(DELPHI		REFID=47716
ABREU	00Z	EPJ C17 53	P. Abreu <i>et al.</i>	(DELPHI	Collab.)	REFID=47790
CASSISI	00	PL B481 323	S. Cassisi <i>et al.</i>			REFID=47907
CHANG	00B	PRL 85 3765	L.N. Chang et al.			REFID=47823
CHEUNG	00	PR D61 015005	K. Cheung			REFID=47885
CORNET	00	PR D61 037701	F. Cornet, M. Relano, J. Rico			REFID=47888
GRAESSER			M.L. Graesser			REFID=47889
	00	PR D61 074019				
HAN	00	PR D62 125018	T. Han, D. Marfatia, RJ. Zhang			REFID=47896
MATHEWS	00	JHEP 0007 008	P. Mathews, S. Raychaudhuri, K. Srid	Ihar		REFID=49019
MELE	00	PR D61 117901	S. Mele, E. Sanchez			REFID=47892
RIZZO	00	PR D61 016007	T.G. Rizzo, J.D. Wells			REFID=47887
ABBIENDI	99P	PL B465 303	G. Abbiendi <i>et al.</i>	(OPAL	Collab.)	REFID=47290
ACCIARRI	99M	PL B464 135	M. Acciarri <i>et al.</i>		Collab.)	REFID=47229
ACCIARRI	99R	PL B470 268	M. Acciarri et al.		Collab.)	REFID=47322
	995				Collab.)	REFID=47323
ACCIARRI		PL B470 281	M. Acciarri <i>et al.</i>	(L3	Collab.)	
BOURILKOV	99	JHEP 9908 006	D. Bourilkov			REFID=47332
HOSKINS	85	PR D32 3084	J.K. Hoskins et al.			REFID=49700